Quantum Joule-Thomson Effect in a Saturated Homogeneous Bose Gas

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We study the thermodynamics of Bose-Einstein condensation in a weakly interacting quasihomogeneous atomic gas, prepared in an optical-box trap. We characterize the critical point for condensation and observe saturation of the thermal component in a partially condensed cloud, in agreement with Einstein's textbook picture of a purely statistical phase transition. Finally, we observe the quantum Joule-Thomson effect, namely isoenthalpic cooling of an (essentially) ideal gas. In our experiments this cooling occurs spontaneously, due to energy-independent collisions with the background gas in the vacuum chamber. We extract a Joule-Thomson coefficient $\mu_{\rm JT} > 10^9$ K/bar, about 10 orders of magnitude larger than observed in classical gases.

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Ultracold atomic gases provide textbooklike demonstrations of basic quantum-statistical phenomena, such as Bose-Einstein condensation [1] and Fermi pressure [2], and allow controllable studies of open problems in the physics of strongly correlated interacting systems [3]. However, they are traditionally produced in harmonic traps, which makes them spatially inhomogeneous and thus different both from other correlated systems, that they often aim to emulate, and from textbook models. This difference can often be addressed by using the local density approximation (LDA) and is in some cases even beneficial [4–10]. However, spatial inhomogeneity is also often problematic. Integrating over the varying density can smear or even qualitatively change experimental signatures (see, e.g., [11–15]). Moreover, LDA breaks down close to phase transitions, where the correlation length diverges [16]. Recently, a Bose-Einstein condensate (BEC) of atoms in an essentially uniform potential was achieved [17], opening new possibilities for studies of both quantum-statistical and interaction effects in many-body systems.

In this Letter, we explore the thermodynamics of a weakly interacting quasihomogeneous Bose gas, prepared in an optical-box trap (see Fig. 1). We characterize the critical point for condensation and demonstrate two purely quantum-statistical phenomena that are, despite weak interactions, obscured in harmonically trapped gases. First, we observe saturation of the thermal component in a partially condensed gas, in agreement with Einstein's textbook picture of condensation as a purely statistical phase transition. Second, we observe and quantitatively explain cooling of partially condensed clouds through isoenthalpic rarefaction, driven by collisions with the background gas in the vacuum chamber. This phenomenon is the quantum version of the Joule-Thomson (JT) effect, exploited in thermal machines such as refrigerators and heat pumps. In the classical JT process, isoenthalpic cooling occurs only due to interactions, whereas a quantum Bose gas is expected to show it even in their absence. This phenomenon was predicted as early as 1937 [18] and was at the time proposed as a way to experimentally observe the quantum statistics obeyed by a gas, the effect being opposite for fermions (see also [19]). However, to our knowledge, JT cooling has never before been experimentally observed in the regime where quantum-statistical effects dominate over the interaction ones. Outside the realm of ultracold atomic gases, the quantum JT effect is many orders of magnitude smaller [18], whereas in ultracold-atom experiments it is usually suppressed because of harmonic trapping and the resulting nonsaturation of the thermal component.

Our experimental setup is described in detail in Refs. [17,20]. We produce BECs of ⁸⁷Rb atoms in the quasiuniform potential of the optical-box trap depicted in Fig. 1. The repulsive walls of the cylindrically shaped box are created by one tubelike and two sheetlike 532-nm laser beams, and we use a magnetic field gradient to cancel the linear gravitational potential. The box trap is loaded from a cloud precooled to $T \sim 100$ nK in a harmonic trap. Further cooling is achieved by forced evaporation in the box potential.

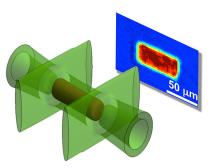


FIG. 1 (color online). Quasihomogeneous Bose gas in an optical-box trap. We show a schematic of our cylindrically shaped trap, formed by intersecting one tubelike and two sheetlike green laser beams, and an *in situ* absorption image of the atomic cloud.

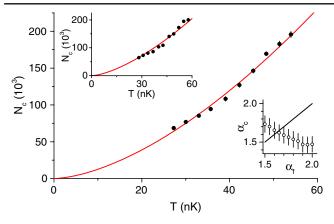


FIG. 2 (color online). Critical point for condensation. A powerlaw fit, $N_c \propto T^{\alpha_c}$, gives $\alpha_c \approx \alpha_f = 1.65$ (see the text). Bottom inset: Within errors, $\alpha_c = \alpha_f$ (solid line) is satisfied for $1.6 \lesssim \alpha_f \lesssim 1.7$. Top inset: Comparison with theory of a perfectly homogeneous gas, with fixed $\alpha_c = \alpha_f = 3/2$.

In a box trap, the BEC spreads over the same volume as the thermal cloud, making the condensate density and the total interaction energy much lower than in harmonically trapped gases with similar atom numbers, temperatures, and *s*-wave scattering length ($a \approx 5$ nm for ⁸⁷Rb). In all our measurements, with condensed fractions up to 40%, the interaction energy is $\lesssim 10\%$ of the thermal energy. We therefore expect the thermodynamics of our clouds to be well described by ideal-gas theory.

We first study the critical atom number for condensation, $N_c(T)$, which also allows us to quantitatively characterize our trap (see Fig. 2). Because of diffraction effects, the optical trap walls cannot be infinitely steep. The resulting small deviation from a perfectly uniform box can to leading order be modeled by an effective high-powerlaw potential, $\propto r^n$ [17]. Defining $\alpha = 3/2 + 3/n$, the thermal momentum distribution is given by the polylog function $g_{\alpha-3/2}$, the density of states is $\rho(\varepsilon) \propto \varepsilon^{\alpha-1}$, and we expect [21]

$$N_c \propto T^{\alpha}$$
. (1)

To measure *T* and the number of thermal (*N*') and condensed (*N*₀) atoms, we image clouds after 50–70 ms of time-of-flight (ToF) expansion from the trap. We vary *T* by changing the evaporation sequence, and the total atom number *N* by changing the initial loading of the trap. After forced evaporation we always raise the trap depth to $U_0 \sim k_B \times 0.4 \,\mu\text{K}$, where evaporation is negligible, and hold the gas for another 1 s before ToF. We identify the critical point with the appearance of small BECs, with $N_0 \lesssim 5000$ [22].

Using Eq. (1) to determine the α value for our trap requires some care, because extracting $N_c(T)$ from ToF images relies on a polylog fitting function that itself

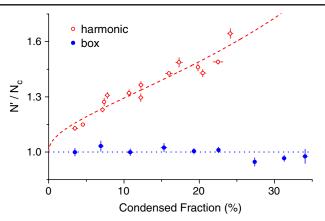


FIG. 3 (color online). Saturation of the thermal component in a partially condensed gas. In the box trap the gas follows the idealgas prediction $N' = N_c$, whereas in the harmonic trap the thermal component is strongly nonsaturated. The dashed red line shows the theoretical prediction for the harmonic trap. (For the harmonic-trap data, $T \approx 110$ nK and $N_c \approx 65 \times 10^3$.)

depends on α . To ensure our analysis is self-consistent, we proceed as follows: We first fit all images using various α values, α_f , and get a separate $N_c(T)$ curve for each α_f . Then, fitting $N_c \propto T^{\alpha_c}$ to these curves as per Eq. (1), we get α_c for each α_f . Finally, requiring $\alpha_c = \alpha_f$ for selfconsistency, we get $1.6 \leq \alpha \leq 1.7$; see the main panel and bottom inset in Fig. 2 [23]. For the data analysis in Figs. 3 and 4 below, we fix $\alpha_f = 1.65$, corresponding to n = 20.

In a perfectly uniform potential, $N_c \propto VT^{3/2}$, where V is the box volume. A small deviation of α from 3/2 corresponds to a slight increase of the effective trapping volume with temperature, $V \propto T^{\alpha-3/2}$, due to the noninfinite steepness of the trap walls. From *in situ* images (as in Fig. 1), we measure a $(20 \pm 10)\%$ increase in V between T = 25 and 50 nK, which is consistent with our α .

Over the whole temperature range in Fig. 2, for the critical phase space density we get $\rho_c = N_c \lambda^3 / V = 2.0 \pm 0.2$, where $\lambda = \hbar \sqrt{2\pi/mk_BT}$ is the thermal wavelength and *m* is the atomic mass. Theoretically, we expect $\rho_c \approx 2.4$, including a finite-size correction [24] to the thermodynamic-limit $\rho_c \approx 2.612$ [21]. This small discrepancy is within our systematic uncertainties of 10% in N_c and 20% in V [25,26].

While in the above we were careful to characterize the small deviation of our trap from an ideal box, we note that simply assuming a perfectly homogeneous gas of constant volume leads to very small errors. This is shown in the top inset in Fig. 2, where we fix both α_f and α_c to 3/2 and the fit of $N_c(T)$ is still very good.

We now turn to the study of partially condensed clouds. In Fig. 3, we show the evolution of the thermal atom number as N is increased beyond N_c and a BEC forms. In Einstein's ideal-gas picture of condensation, N' saturates

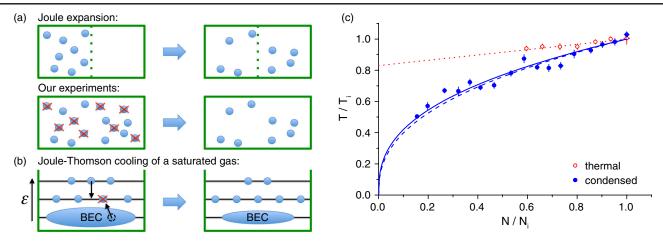


FIG. 4 (color online). Quantum Joule-Thomson effect. (a) Isoenthalpic rarefaction. In a conventional JT process V is increased, whereas in our experiments N is reduced, but for an ideal gas both processes conserve specific enthalpy and are thermodynamically equivalent. (b) Microscopic origin of JT cooling in a saturated quantum gas. Removal of a thermal atom requires a zero-energy particle to come out of the BEC in order to maintain $N' = N_c$, while energy conservation requires redistribution of atoms between energy levels. Cooling is seen in the change of the relative populations of the excited states. In this cartoon we neglect the fact that N_c also slowly decreases. (c) Measurements with a partially condensed and a thermal cloud. The solid and dashed blue lines are predictions of Eq. (2) with $\alpha = 1.65$ and $\alpha = 3/2$, respectively. The dotted red line is a numerical calculation (see the text).

at the critical point and can never exceed N_c . In a harmonically trapped gas, this picture is strongly violated even for weak interactions, such as in ⁸⁷Rb, and can be recovered only by extrapolating to the strictly noninteracting limit [14], where direct measurements cannot be performed due to the absence of thermal equilibrium [28]. This strong deviation from ideal-gas behavior arises due to an interplay of interactions and the nonuniformity of the gas, and can be explained by using a mean-field theory that does not require violation of the saturation picture at the level of the local thermal density [14,29].

In our experiments, N' is essentially equivalent to the thermal density. In Fig. 3, we see that in this case N' indeed remains constant as the condensed fraction N_0/N is increased at constant temperature (45 ± 1 nK) [30]. For comparison, we also show analogous measurements for the harmonic trap from which our box trap is loaded. Here the nonsaturation of N' is prominent and agrees with the prediction based on Refs. [14,29], without any free parameters.

Finally, we study the free long-time evolution of a gas held in the box trap and observe the quantum Joule-Thomson effect (Fig. 4). We prepare clouds at $T_i \approx$ 45 nK in a trap of depth U_0 , where evaporation is negligible. As the gas is held in the trap, atoms are slowly removed from the cloud, with an exponential time scale $\tau \approx 10$ s, by collisions with the background gas in the vacuum chamber. Because of low atomic density ($<5 \times 10^{12}$ cm⁻³), the three-body recombination rate is negligible [34]. As N decays, the elastic collision rate (among the trapped atoms) remains sufficiently high for the gas to continuously reequilibrate on a time scale $\tau_{eq} \lesssim$ 2 s $\ll \tau$ [35]. Crucially, collisions with the background gas are independent of the energy of the trapped atoms, so the average energy per trapped particle, E/N, remains constant. Within the ideal-gas approximation, pressure is simply given by the energy density, so enthalpy H is simply proportional to E, and the specific enthalpy h = H/N is a conserved quantity. The decay of the cloud is thus equivalent to the Joule-Thomson isoenthalpic rarefaction, as depicted in Fig. 4(a). For simplicity, here we illustrate free Joule expansion into vacuum; in the JT process rarefaction is driven by throttling, but for an ideal gas both processes are isoenthalpic.

In a *classical* ideal gas, *h* depends only on *T*, so temperature cannot change in an isoenthalpic process. This is, however, not true for a quantum degenerate ideal gas [18,21]. In a partially condensed cloud, only thermal atoms contribute to the total energy, so for a saturated gas $E \propto N'T = N_c T$. More precisely, $E = \alpha [\zeta(\alpha + 1)/\zeta(\alpha)]N_c k_B T$ [21], where ζ is the Riemann zeta function. From Eq. (1), keeping E/N constant implies cooling according to

$$T \propto N^{1/(\alpha+1)}.$$
 (2)

The microscopic origin of this cooling is qualitatively illustrated in Fig. 4(b). Removal of any thermal atom requires a zero-energy particle to come out of the BEC in order to maintain the saturation of the thermal component. Energy conservation then requires simultaneous redistribution of atoms between energy levels. The net cooling is seen in the change of the relative populations of the excited states. Note that direct removal of BEC atoms does not change the temperature and also that in this simplified cartoon we neglect the fact that $N_c(T)$ gradually decreases as T drops.

In Fig. 4(c), we show the evolution of T with decaying N, with both quantities scaled to their initial values. We show data for a partially condensed gas, with $N_i \approx 1.7N_c(T_i)$, and a thermal cloud, with $N_i \approx 0.6N_c(T_i)$. Note that in both cases we measure down to approximately the same final N value, $\approx 0.3N_c(T_i)$. At this point, N_0 in the initially condensed sample vanishes, and also below this N our temperature fits become less reliable.

In the (partially) condensed sample we observe a drop in *T* by a factor of 2, i.e., by ≈ 22 nK. Meanwhile, the interaction energy per particle is always smaller than $(8\pi\hbar^2 a/m)N_i/V$ [21], corresponding to < 4 nK. The change in *T* therefore must predominantly be a purely quantum-statistical, rather than an interaction, effect [36].

Indeed, Eq. (2) fits the data very well. We show predictions for $\alpha = 1.65$ (solid blue line) and $\alpha = 3/2$ (dashed blue line). The two are almost indistinguishable, reaffirming that the behavior of our clouds is very close to that of a perfectly homogeneous gas.

Within the constant-*V* approximation, our measurements are directly related to the Joule-Thomson coefficient $\mu_{JT} = (\partial T / \partial P)_h$, according to $\mu_{JT} = A(\partial T / \partial N)_h$, with constant A = (3/2)VN/E. More conveniently expressed, the dashed blue line in Fig. 4(c) corresponds to [21]

$$\mu_{\rm JT} = \frac{2}{5\zeta(5/2)} \frac{\lambda^3}{k_B}.$$
(3)

At our lowest temperatures, $\mu_{JT} \approx 4 \times 10^9$ K/bar, about 10 orders of magnitude larger than observed in classical gases.

To understand the origins of this enhancement, on dimensional grounds we write $\mu_{JT} = c_{JT}T/P$, so c_{JT} is dimensionless and $T/P \sim 1/(n'k_B)$, where n' is the thermal particle density. The value of $c_{\rm JT}$ depends on the equation of state. In the ideal gas, $c_{\rm JT} = 0$ in the classical limit, while $c_{\rm JT} = 1/(\alpha + 1)$ for a saturated Bose gas [21]. Interaction corrections to $c_{\rm JT}$ are essentially given by the ratio of interaction and thermal energy. In our case these small corrections are unimportant and $c_{\rm JT} \approx 2/5$, but for a classical gas they are the only origin of the JT effect and typically $c_{\rm JT} \lesssim$ 10^{-3} [37]. Another large enhancement of $\mu_{\rm IT}$ comes from the fact that in a saturated gas $n' \sim \lambda^{-3} \sim T^{3/2}$ decreases with temperature, and our lowest n' is $\sim 10^7$ times smaller than at ambient temperature and pressure. Combining the differences in $c_{\rm JT}$ and 1/n', we recover the ~10¹⁰-fold increase in $\mu_{\rm IT}$ compared to classical gases. It is also interesting to note that $\mu_{\rm IT}$ in Eq. (3) explicitly vanishes in the limit $\hbar \to 0$, reiterating that this is a purely quantum effect.

Compared to the partially condensed gas, in the thermal sample T remains almost constant, but we do discern a slight cooling effect. This is indeed expected for a degenerate noncondensed gas [18], with a chemical potential

 $\mu \gtrsim -k_B T$. Degeneracy preferentially enhances occupation of low energy states, with $\varepsilon < |\mu|$ (e.g., momentum distribution is more "peaky" than a Gaussian [17,21]), so E/N is lower than the classical value $\alpha k_B T$. Between the classical limit and the critical point, $\eta \equiv E/(\alpha N k_B T)$ decreases from 1 to $\zeta(\alpha + 1)/\zeta(\alpha)$ [21]. Hence, if E/N is kept constant while η grows as the gas becomes less degenerate due to N decay, T must slightly decrease. The dotted red line in Fig. 4(c) shows a numerical calculation of this effect for our $N_i/N_c(T_i)$ and fits the data very well.

Similar theoretical analysis applies to an ideal harmonically trapped gas, but in that case $\alpha = 3$, implying weaker cooling in both condensed clouds and degenerate thermal samples; see Eq. (2) and note that $\zeta(4)/\zeta(3) \approx 0.9$, as compared to $\zeta(5/2)/\zeta(3/2) \approx 0.5$. Moreover, in practice harmonically trapped gases strongly deviate from this picture, because N' is not saturated [14], much higher typical BEC densities enhance the role of interactions, and threebody recombination continuously increases E/N.

In conclusion, we have characterized the critical point for Bose-Einstein condensation in a quasihomogeneous weakly interacting gas and demonstrated two textbooklike quantum-statistical phenomena that highlight qualitative differences between uniform and harmonically trapped degenerate gases. In future work, combining our methods with stronger interactions (employing a Feshbach resonance [38]) and flexible shaping of boxlike traps [39] should allow further studies of novel thermomechanical effects [40].

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- [1] M. H. Anderson, J. R. Ensher, M. R. Matthews, C. E. Wieman, and E. A. Cornell, Science **269**, 198 (1995).
- [2] A. G. Truscott, K. E. Strecker, W. I. McAlexander, G. B. Partridge, and R. G. Hulet, Science 291, 2570 (2001).
- [3] I. Bloch, J. Dalibard, and W. Zwerger, Rev. Mod. Phys. 80, 885 (2008).
- [4] T.-L. Ho and Q. Zhou, Nat. Phys. 6, 131 (2010).
- [5] S. Nascimbène, N. Navon, K. J. Jiang, F. Chevy, and C. Salomon, Nature (London) 463, 1057 (2010).
- [6] N. Navon, S. Nascimbène, F. Chevy, and C. Salomon, Science 328, 729 (2010).
- [7] C.-L. Hung, X. Zhang, N. Gemelke, and C. Chin, Nature (London) 470, 236 (2011).
- [8] T. Yefsah, R. Desbuquois, L. Chomaz, K. J. Günter, and J. Dalibard, Phys. Rev. Lett. 107, 130401 (2011).
- [9] R. P. Smith, N. Tammuz, R. L. D. Campbell, M. Holzmann, and Z. Hadzibabic, Phys. Rev. Lett. **107**, 190403 (2011).

- [10] M. J. H. Ku, A. T. Sommer, L. W. Cheuk, and M. W. Zwierlein, Science 335, 563 (2012).
- [11] C. Raman, M. Köhl, R. Onofrio, D. S. Durfee, C. E. Kuklewicz, Z. Hadzibabic, and W. Ketterle, Phys. Rev. Lett. 83, 2502 (1999).
- [12] A. Perali, F. Palestini, P. Pieri, G. C. Strinati, J. T. Stewart, J. P. Gaebler, T. E. Drake, and D. S. Jin, Phys. Rev. Lett. 106, 060402 (2011).
- [13] S. Nascimbène, N. Navon, S. Pilati, F. Chevy, S. Giorgini, A. Georges, and C. Salomon, Phys. Rev. Lett. 106, 215303 (2011).
- [14] N. Tammuz, R. P. Smith, R. L. D. Campbell, S. Beattie, S. Moulder, J. Dalibard, and Z. Hadzibabic, Phys. Rev. Lett. 106, 230401 (2011).
- [15] R. P. Smith, R. L. D. Campbell, N. Tammuz, and Z. Hadzibabic, Phys. Rev. Lett. **106**, 250403 (2011).
- [16] T. Donner, S. Ritter, T. Bourdel, A. Öttl, M. Köhl, and T. Esslinger, Science 315, 1556 (2007).
- [17] A. L. Gaunt, T. F. Schmidutz, I. Gotlibovych, R. P. Smith, and Z. Hadzibabic, Phys. Rev. Lett. 110, 200406 (2013).
- [18] D. S. Kothari and B. N. Srivasava, Nature (London) 140, 970 (1937).
- [19] E. Timmermans, Phys. Rev. Lett. 87, 240403 (2001).
- [20] I. Gotlibovych et al., arXiv:1212.4108.
- [21] C. Pethick and H. Smith, *Bose-Einstein Condensation in Dilute Gases* (Cambridge University Press, Cambridge, England, 2002).
- [22] Since N' saturates for $N > N_c$, as shown in Fig. 3, associating the critical point with a small but nonzero N_0 does not introduce any errors.
- [23] This conclusion is also compatible with an earlier estimate based on analyzing just the shape of the thermal ToF profiles at \sim 100 nK [17].
- [24] S. Grossmann and M. Holthaus, Z. Phys. B 97, 319 (1995).
- [25] Our absolute atom-number calibration is obtained by measuring N_c in a harmonic trap [15]. The systematic uncertainty in V is due to the inherent difficulties in imaging small and dense clouds. Note that it corresponds to about one CCD pixel uncertainty in the linear cloud size.

- [26] N_c could also be slightly reduced by weak disorder in the optical potential [27]. By directly imaging our trapping beams, we assess the disorder in our trap to be up to a few percent of U_0 .
- [27] T. Bourdel, Phys. Rev. A 86, 063626 (2012).
- [28] A. L. Gaunt, R. J. Fletcher, R. P. Smith, and Z. Hadzibabic, Nat. Phys. 9, 271 (2013).
- [29] R. P. Smith and Z. Hadzibabic, in *Physics of Quantum Fluids*, Springer Series in Solid-State Sciences Vol. 177, edited by A. Bramati and M. Modugno (Springer, Berlin, 2013), Chap. 16, pp. 341–359; see also arXiv:1203.2063.
- [30] In an interacting gas the thermal phase-space density can actually drop below ρ_c once the BEC forms [31–33], but for our parameters this effect is very small.
- [31] V. V. Goldman, I. F. Silvera, and A. J. Leggett, Phys. Rev. B 24, 2870 (1981).
- [32] D. A. Huse and E. D. Siggia, J. Low Temp. Phys. 46, 137 (1982).
- [33] A. Minguzzi, S. Conti, and M. P. Tosi, J. Phys. Condens. Matter 9, L33 (1997).
- [34] J. Söding, D. Guéry-Odelin, P. Desbiolles, F. Chevy, H. Inamori, and J. Dalibard, Appl. Phys. B 69, 257 (1999).
- [35] C. R. Monroe, E. A. Cornell, C. A. Sackett, C. J. Myatt, and C. E. Wieman, Phys. Rev. Lett. **70**, 414 (1993).
- [36] We estimate that the total change in *T* due to the change in the interaction energy per particle could at most be \sim 3 nK, which is comparable to our experimental data scatter. Moreover, note that for *a* > 0 the drop in the interaction energy should lead to slight heating rather than cooling of the gas.
- [37] For a classical gas, $c_{\rm JT}$ can be estimated, to within an order of magnitude, by using the van der Waals approximation for the equation of state.
- [38] C. Chin, R. Grimm, P. Julienne, and E. Tiesinga, Rev. Mod. Phys. 82, 1225 (2010).
- [39] A. L. Gaunt and Z. Hadzibabic, Sci. Rep. 2, 721 (2012).
- [40] D. J. Papoular, G. Ferrari, L. P. Pitaevskii, and S. Stringari, Phys. Rev. Lett. **109**, 084501 (2012).